

Searching for sources of carbonaceous dust grains with JWST spectroscopy

Dusty Universe 2025

University of Arizona November 11th, 2025

Sam Rose (Caltech)

With Mansi Kasliwal (Caltech), Ryan Lau
(IPAC), Jacob Jencson (IPAC), Viraj
Karambelkar (Columbia), Howard Bond
(STScI/Penn State), and the WR-Dust-ERS
team

Sam Rose (Caltech) srose@caltech.edu

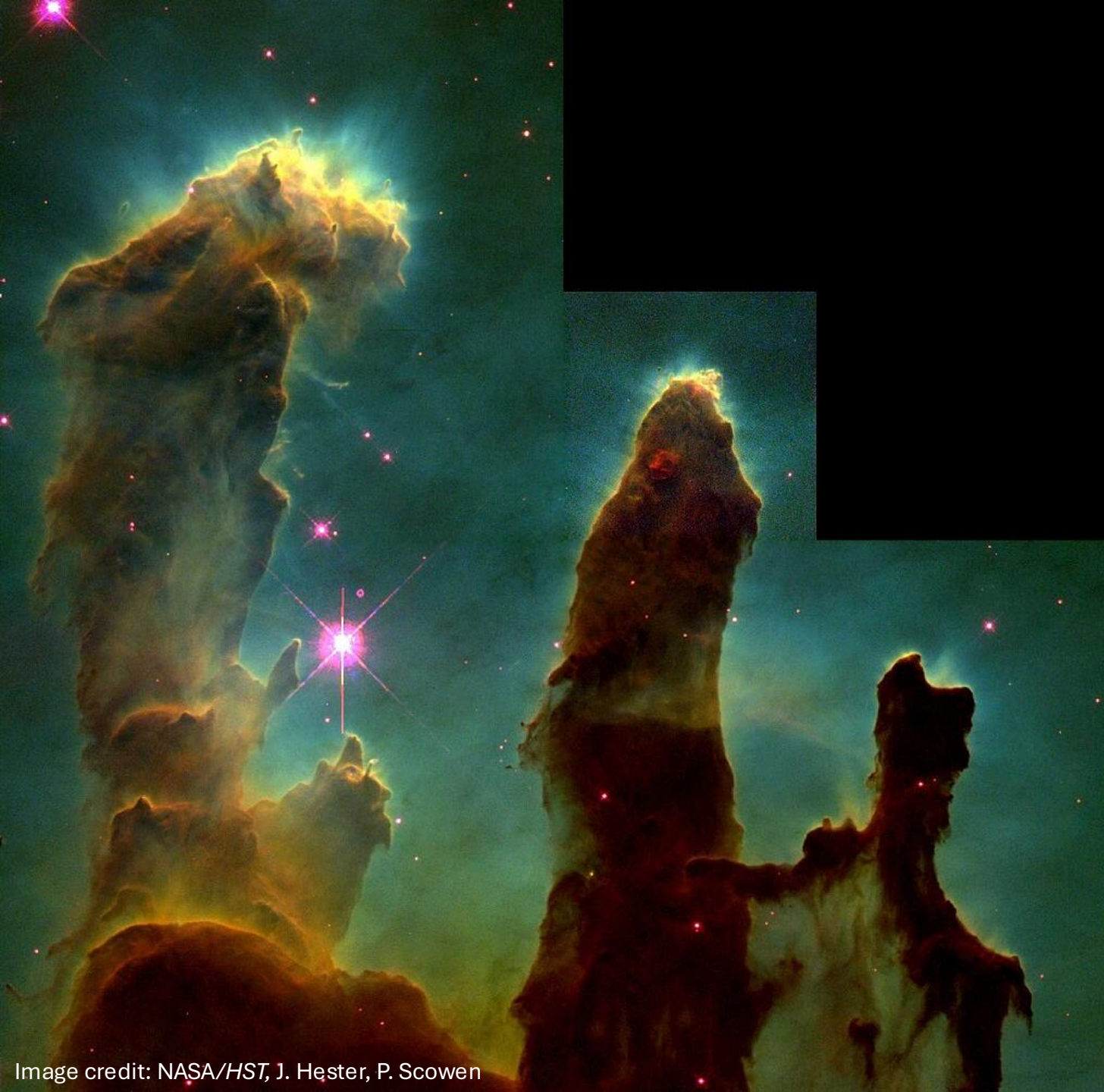
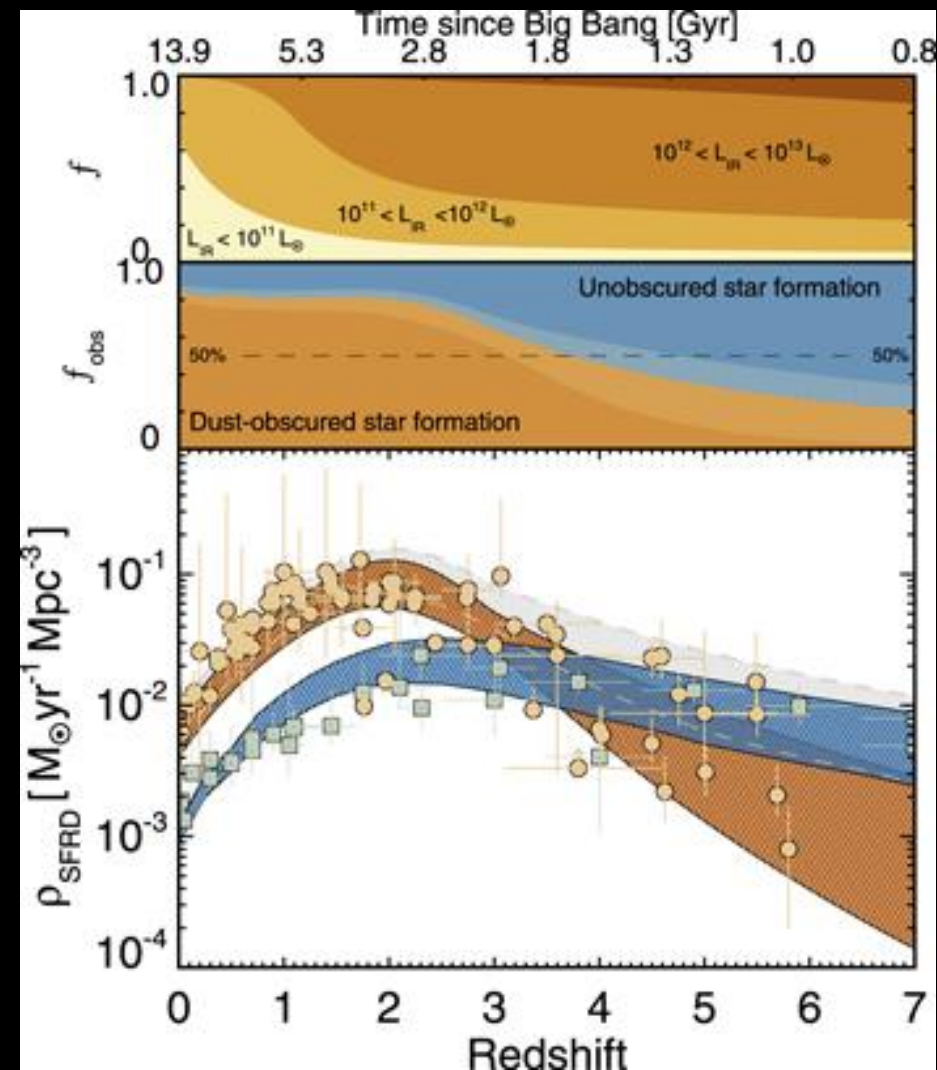
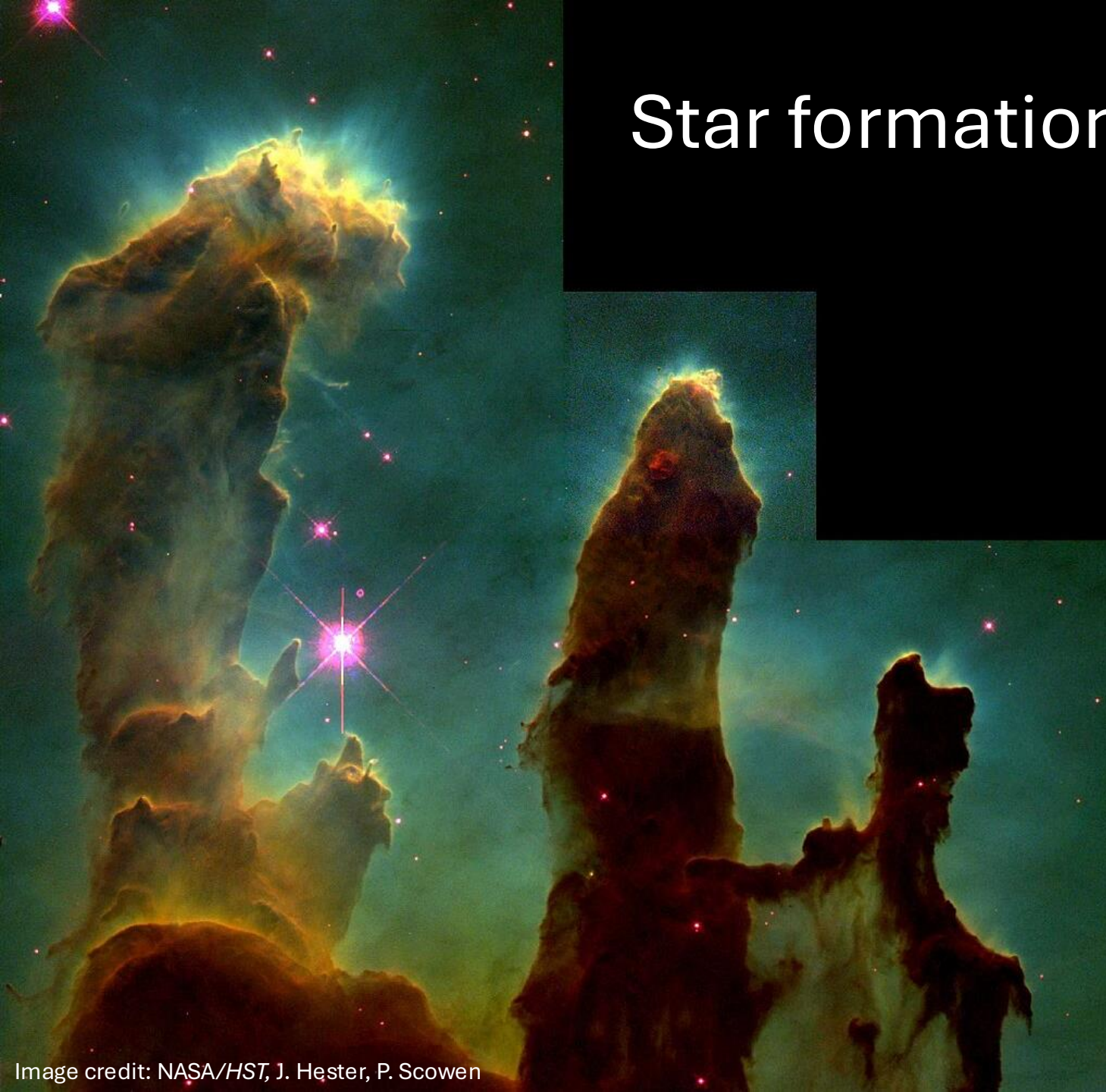


Image credit: NASA/HST, J. Hester, P. Scowen

Why study astronomical dust formation?

Star formation is dust obscured!



Zavala+ 2021

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srose@caltech.edu

Dust rocks!



Image credit: NASA/Apollo 17

Dust chemistry (silicate
versus carbonaceous)
impacts planet
formation!



Image credit: NASA

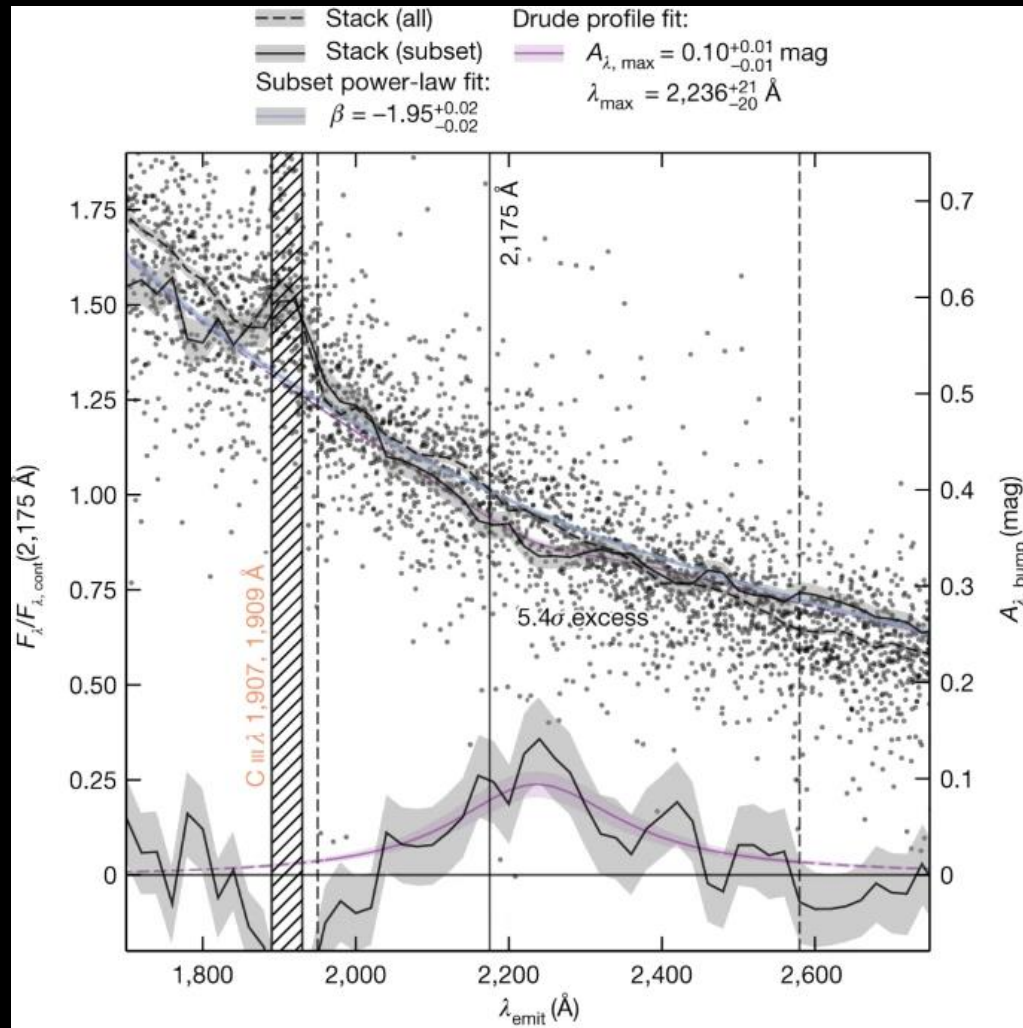


Image credit: NASA

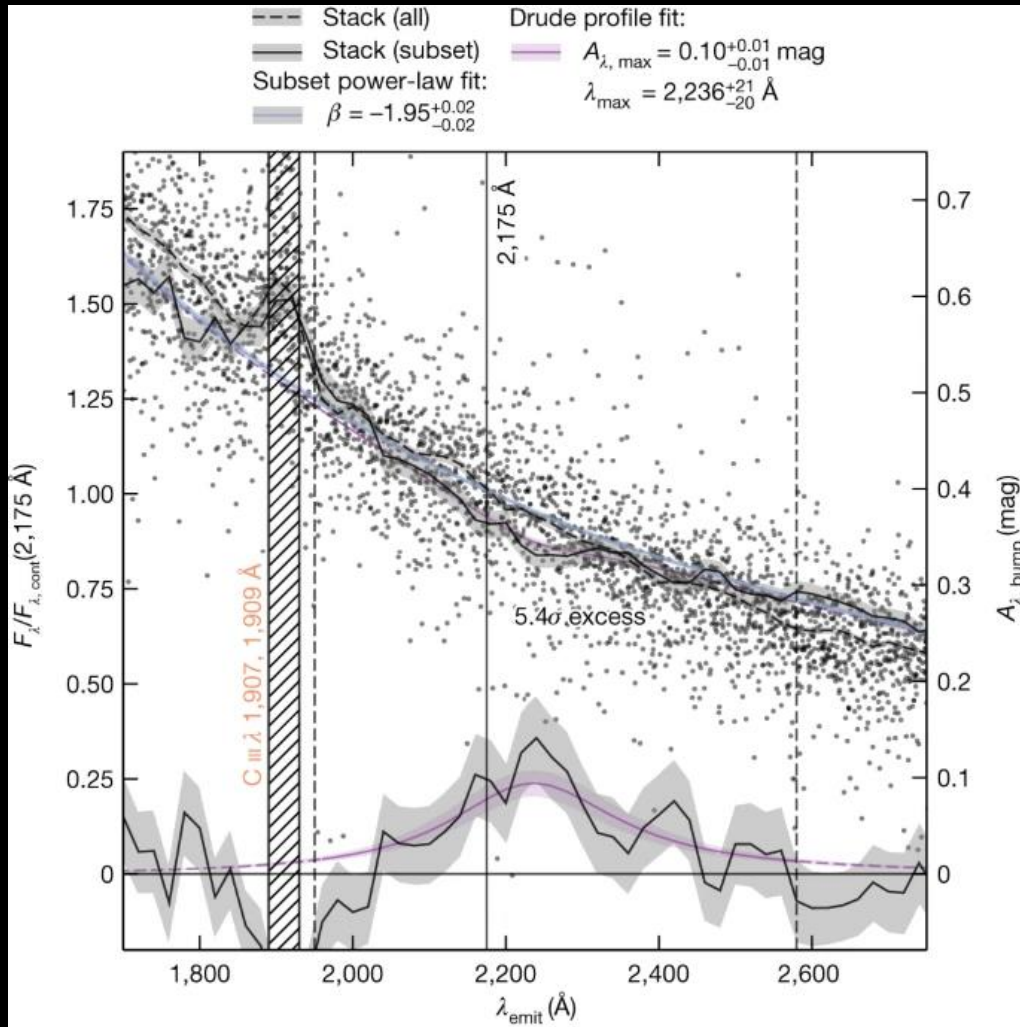


Image credit: S. Rose

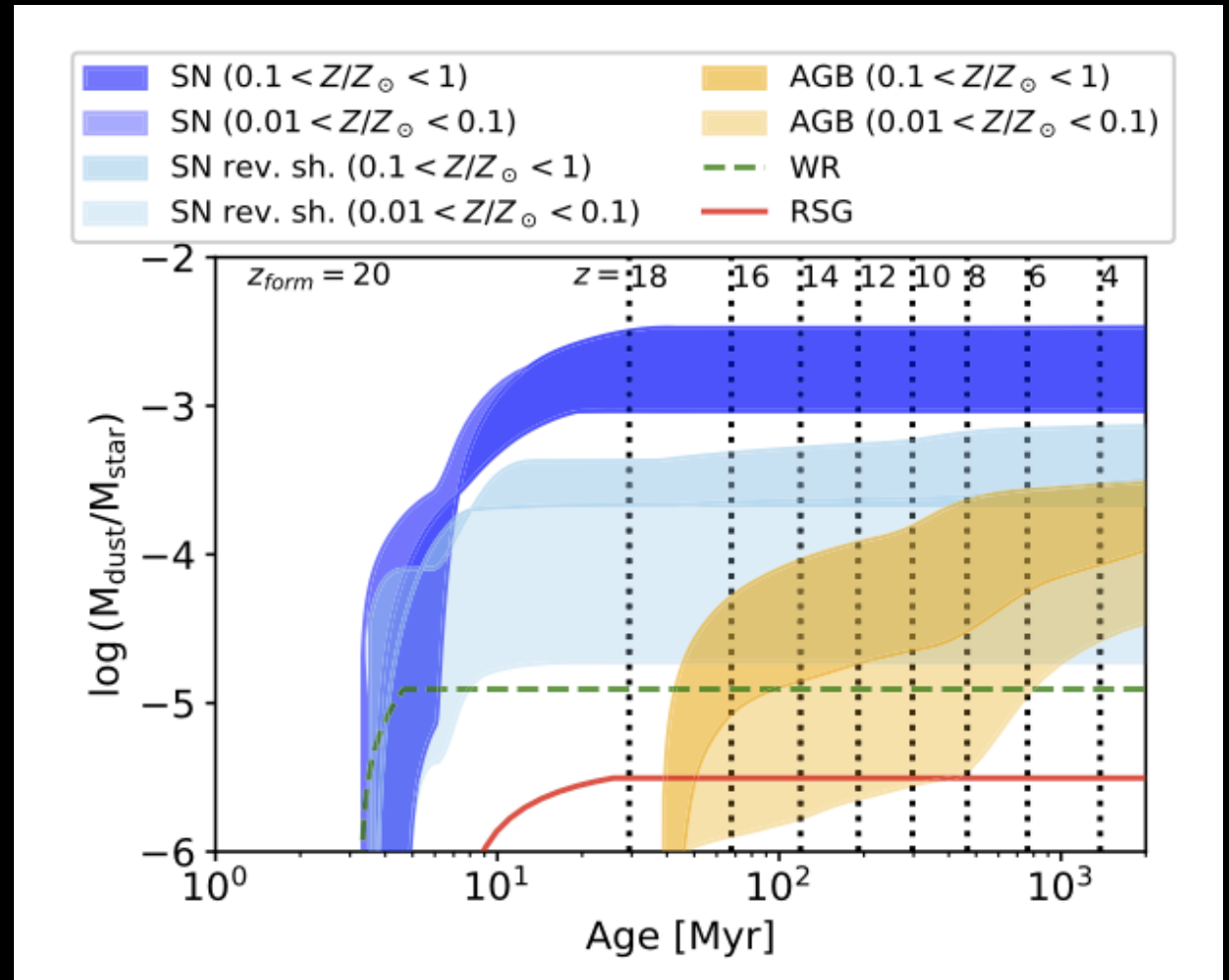
Dust forms early in the Universe – including carbonaceous dust grains!



Supernovae form dust early



Witstok+ 2023

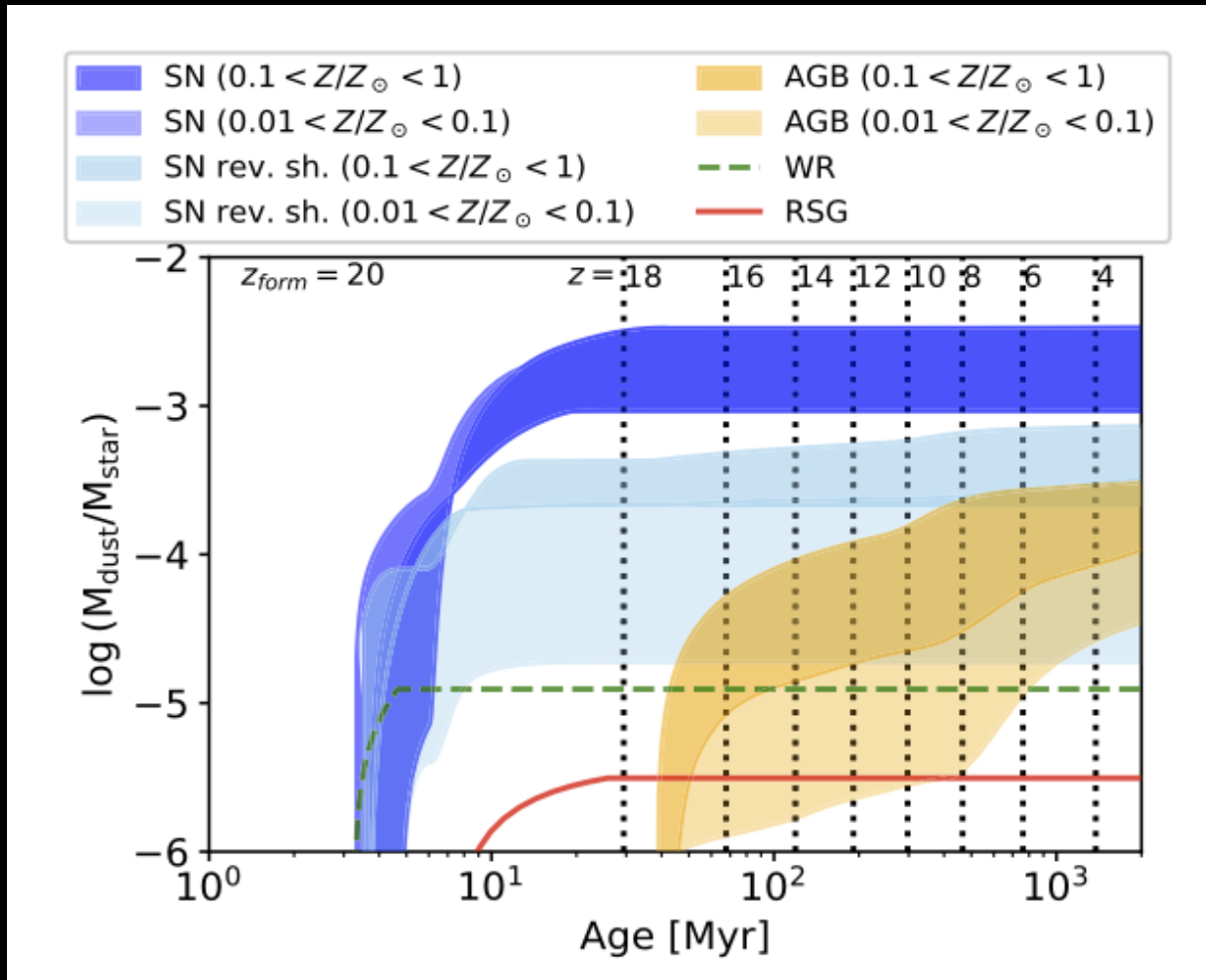


Schneider and Maiolino 2024

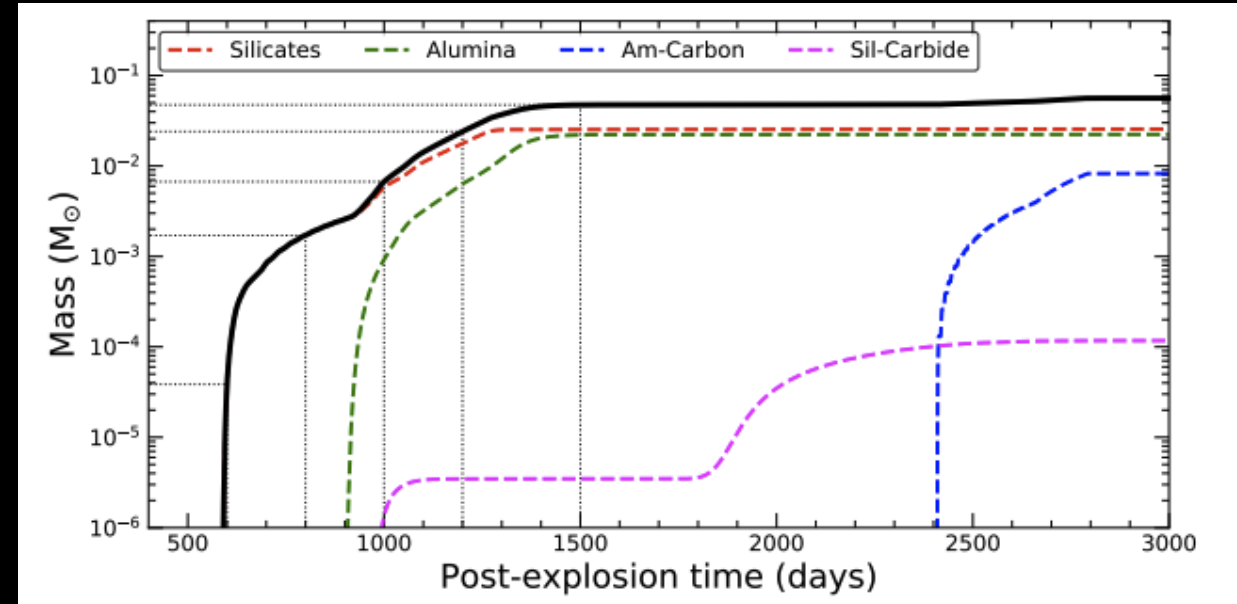
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srose@caltech.edu

... but primarily form O-rich silicate dust



Schneider and Maiolino 2024



Sarangi 2022

See also: Gall+ 2014, Shahbandeh+2023,2025, Tinyanont +2025

Also, what about dust destruction by the reverse shock? E.g., (Micelotta+2018)

AGB stars form C-rich dust ... but not early enough

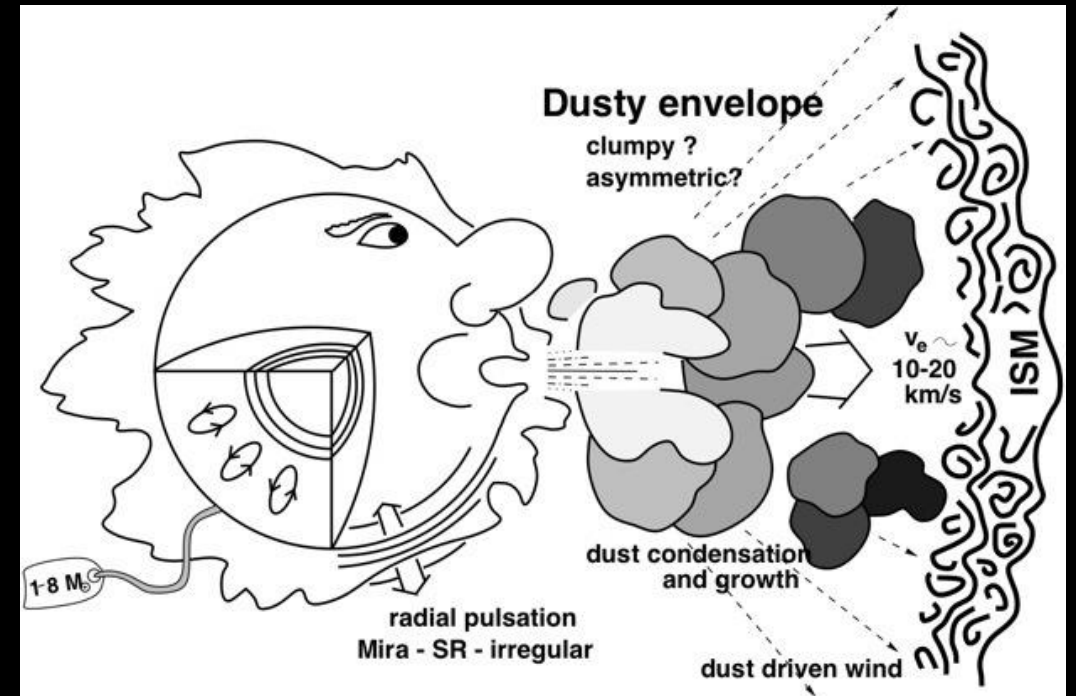
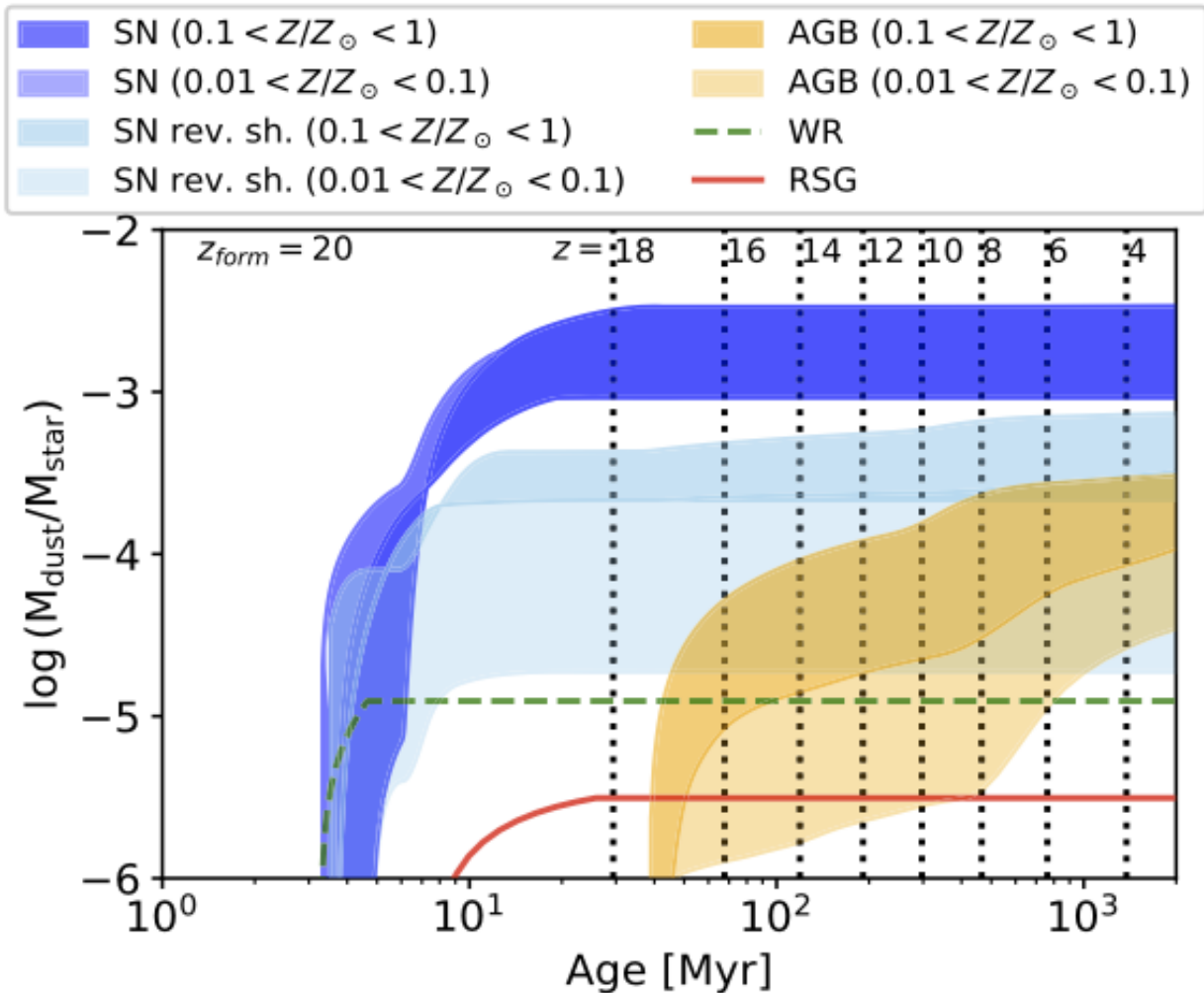
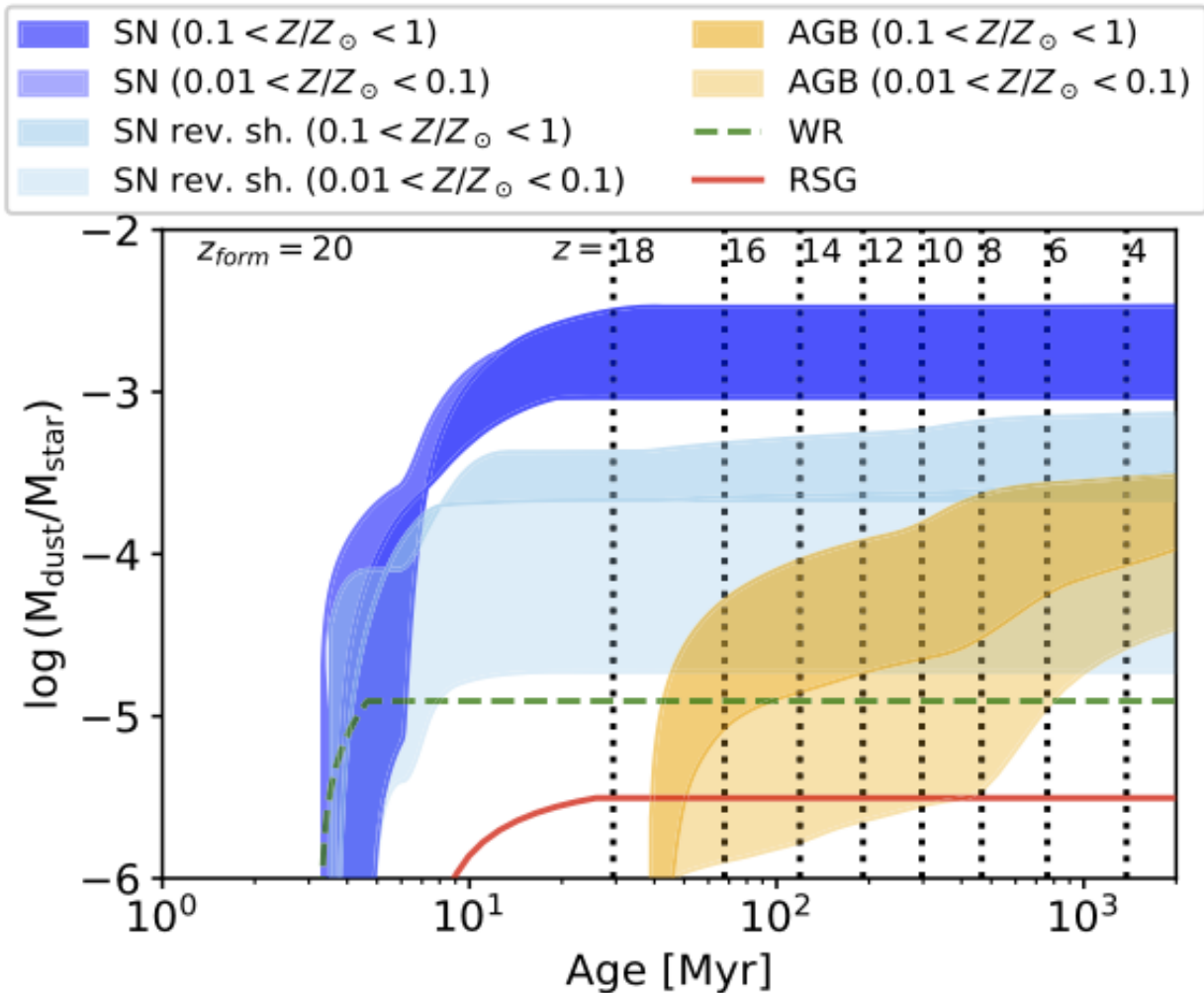


Image credit: M. Marengo CfA

Are there other important sources of dust?



WR 140: *JWST*/NASA/ESA Lau +2022

Riko Senoo's earlier talk on WR 140!

Can you get carbon-rich dust
producing WR stars at low
metallicity?

How do you get a dust producing WR system?

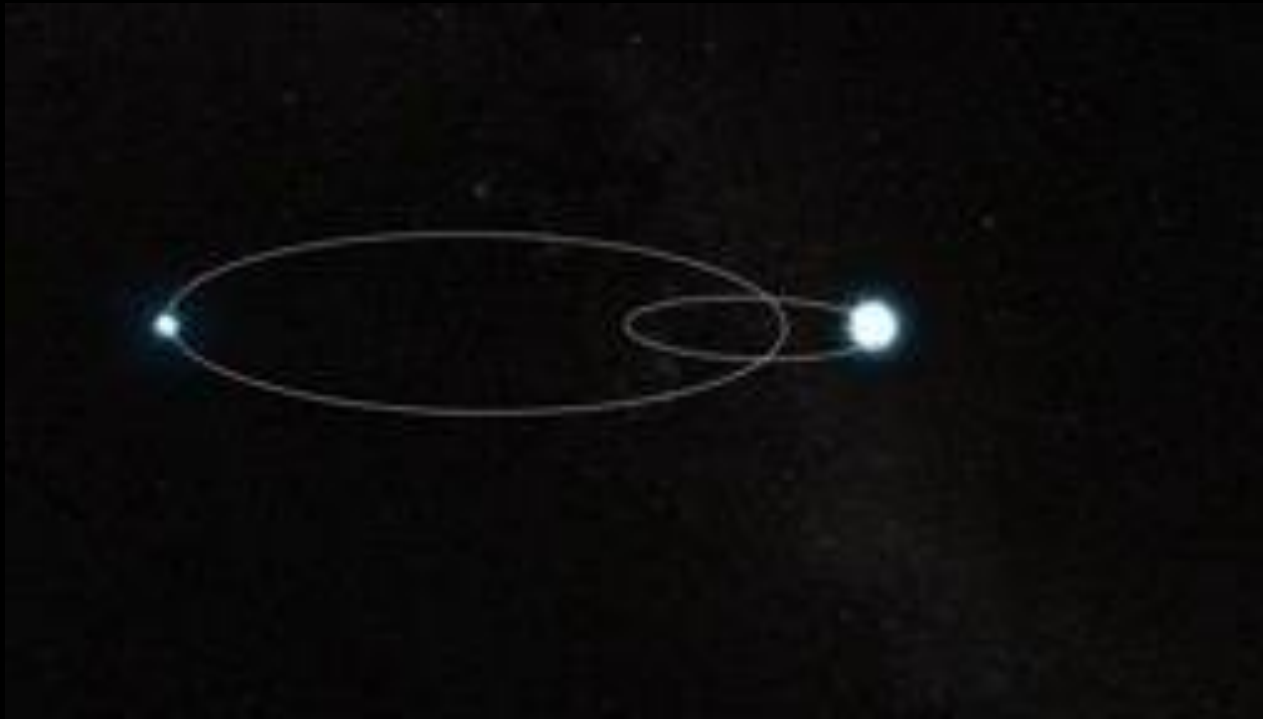


- WR stars form from massive evolved stars
- Hot (20,000-200,000 K) exposed C/O (N) core
- Characterized by broad features
- High wind velocities + mass loss rates

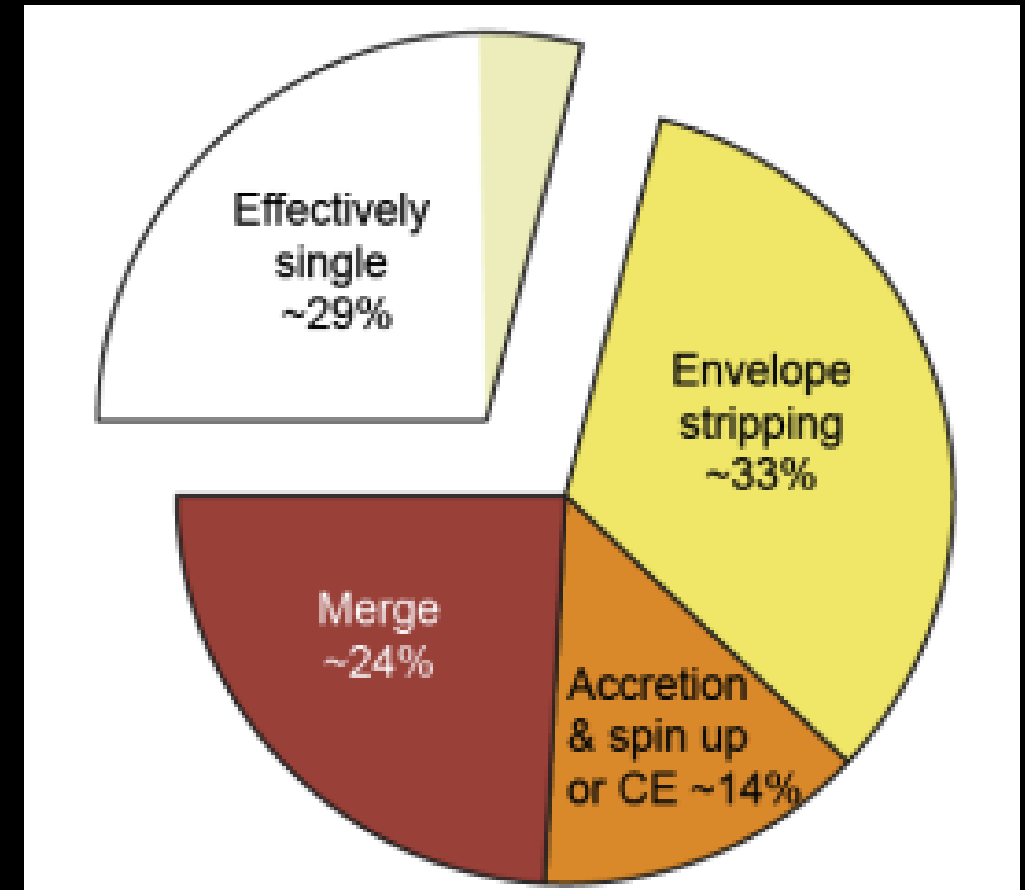
A little surprising that we get dust formation in such an inhospitable environment!

See, e.g., Allen+1972, Williams+1987,2001,2009, Usov+1991,Lau+2020a

How do you get a dust producing WR system?

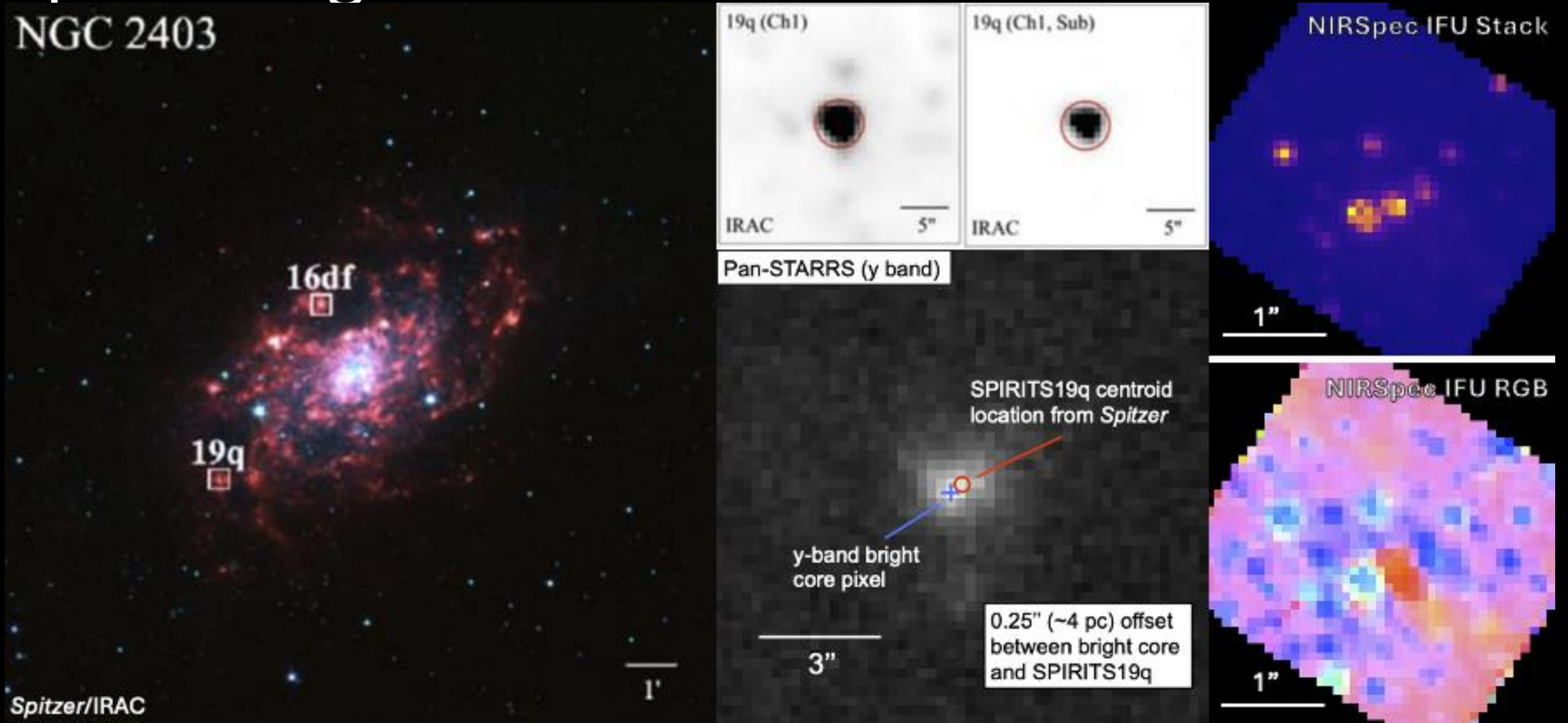


See e.g., Allen+1972, Williams+1987, Usov+1991, Lau+2020a



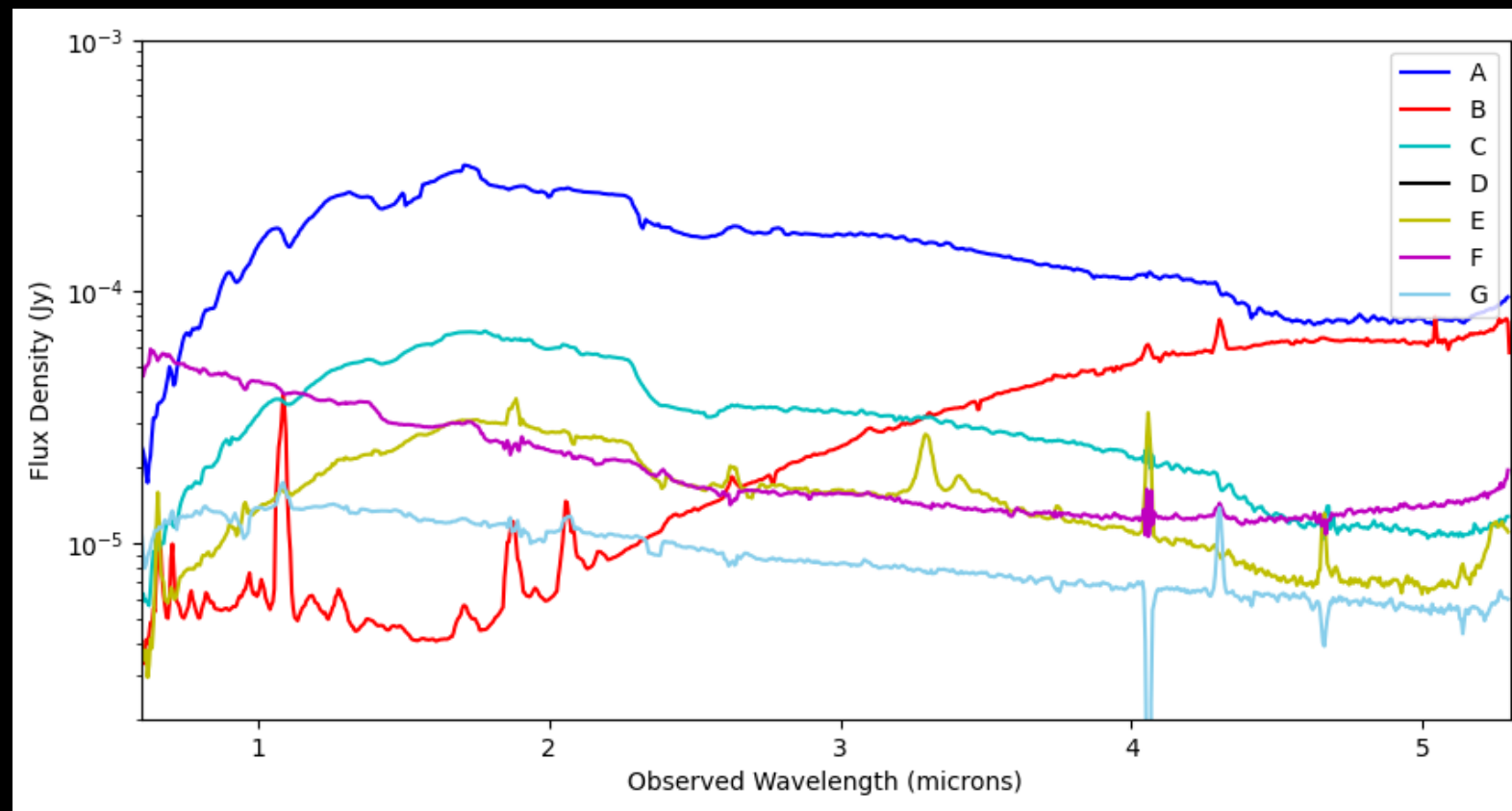
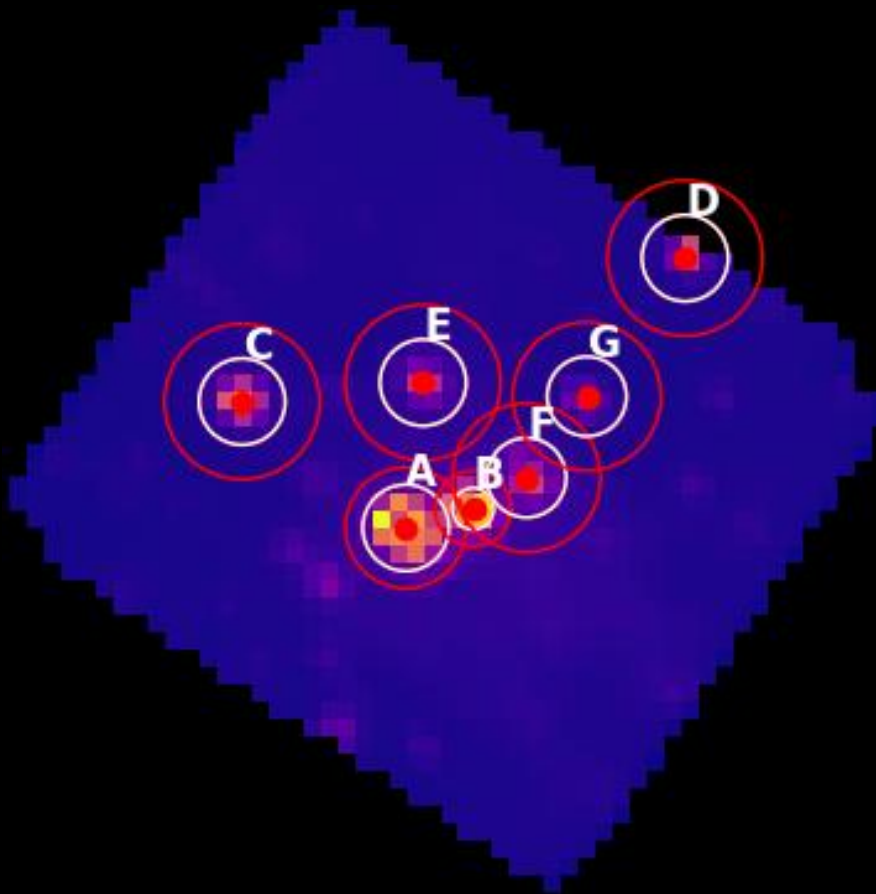
Sana+ 2012, see also Dsilva+2020

SPIRITS 19q: a low metallicity highly dust producing WC candidate



SPIRITS 19q is a WC star!

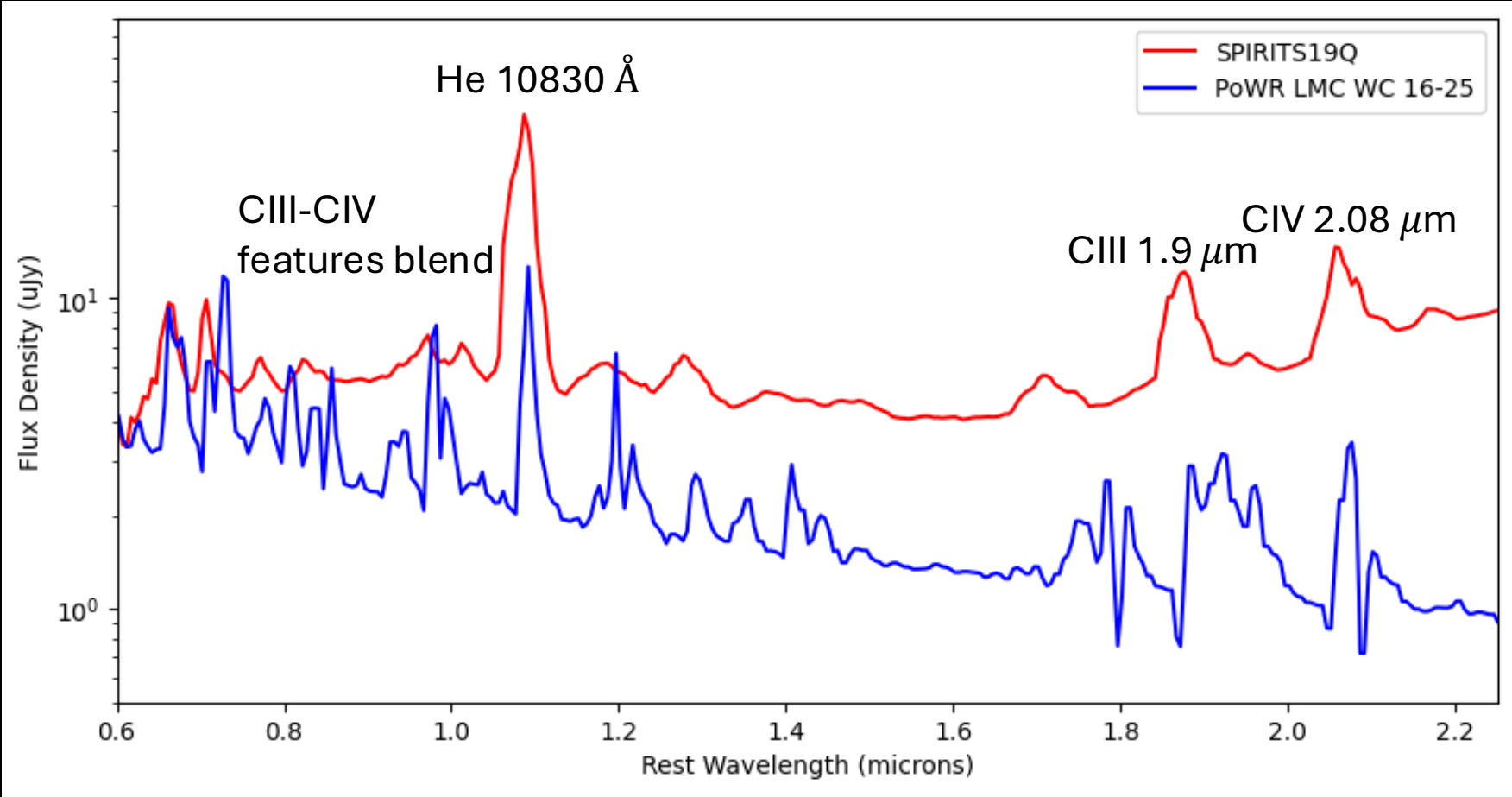
- SPIRITS 19q is red: $F_{4.5\mu\text{m}} > F_{3.6\mu\text{m}}$



SPIRITS 19q is a highly dust producing WC star!

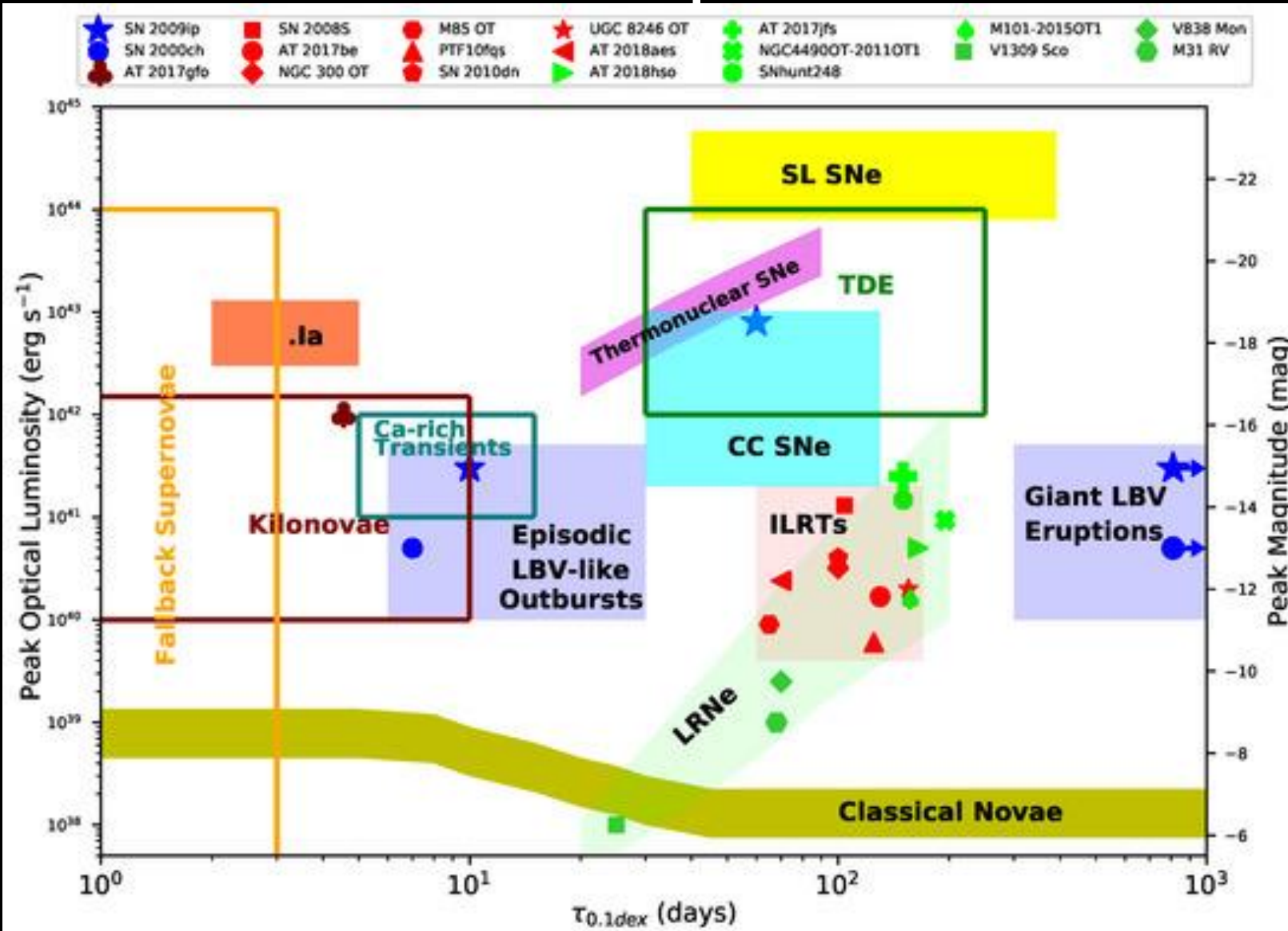
Up to $4 \times 10^{-6} M_{\odot}/\text{yr}$ of dust forming (Lau+2021)!
All AGB stars in the LMC: $1.4 \times 10^{-5} M_{\odot}/\text{yr}$ (Riebel+2012)

$\text{Log}(T_{\text{eff}}/\text{K})$	5.15
$\text{Log}(L/L_{\odot})$	5.3
$R_{\star}$	$0.75 R_{\odot}$
$\text{Log}(\dot{M}/\frac{M_{\odot}}{\text{yr}})$	-4.2
V_f	2000 km/s
$M_{\star}$	$11.64 M_{\odot}$



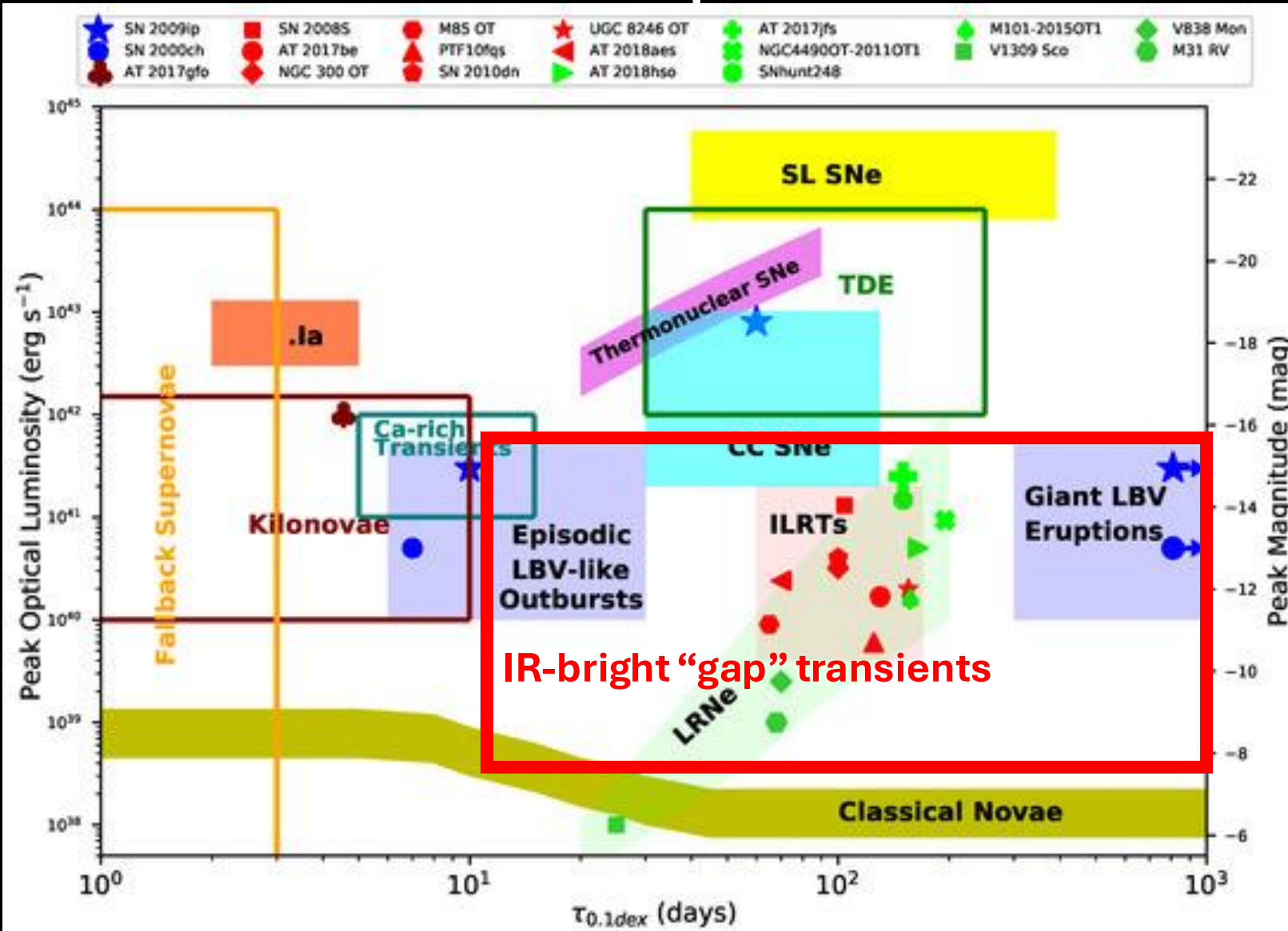
There is at least one highly dust producing WC
star system at LMC metallicity! We should
look for more...

Are there other important sources of dust?



A whole landscape of astronomical transients with lots of mass ejected!

Are there other important sources of dust?



A whole landscape of astronomical transients with lots of mass ejected!

Are there dust producing
transients in the luminosity gap
and do they make
carbonaceous dust?

JWST Cycle 4 7040: A MIR Survey of “gap” transients

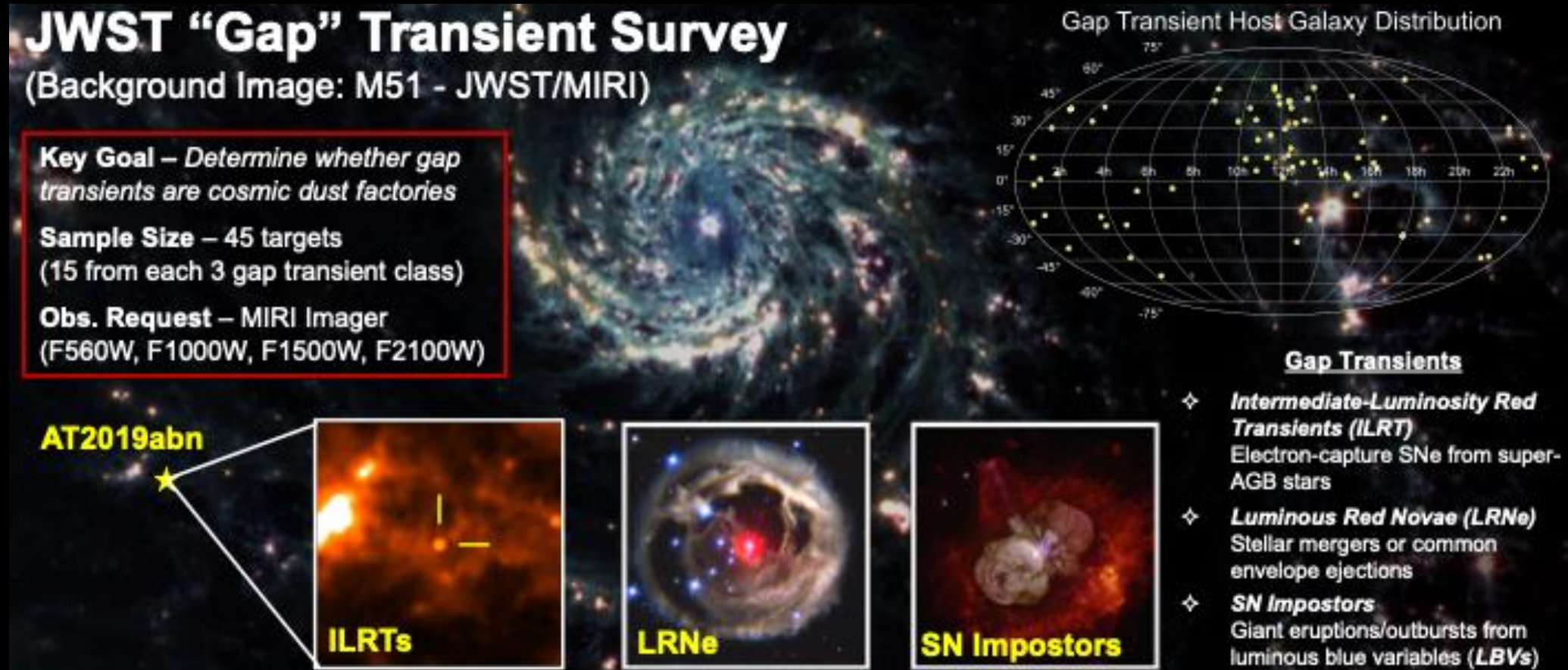
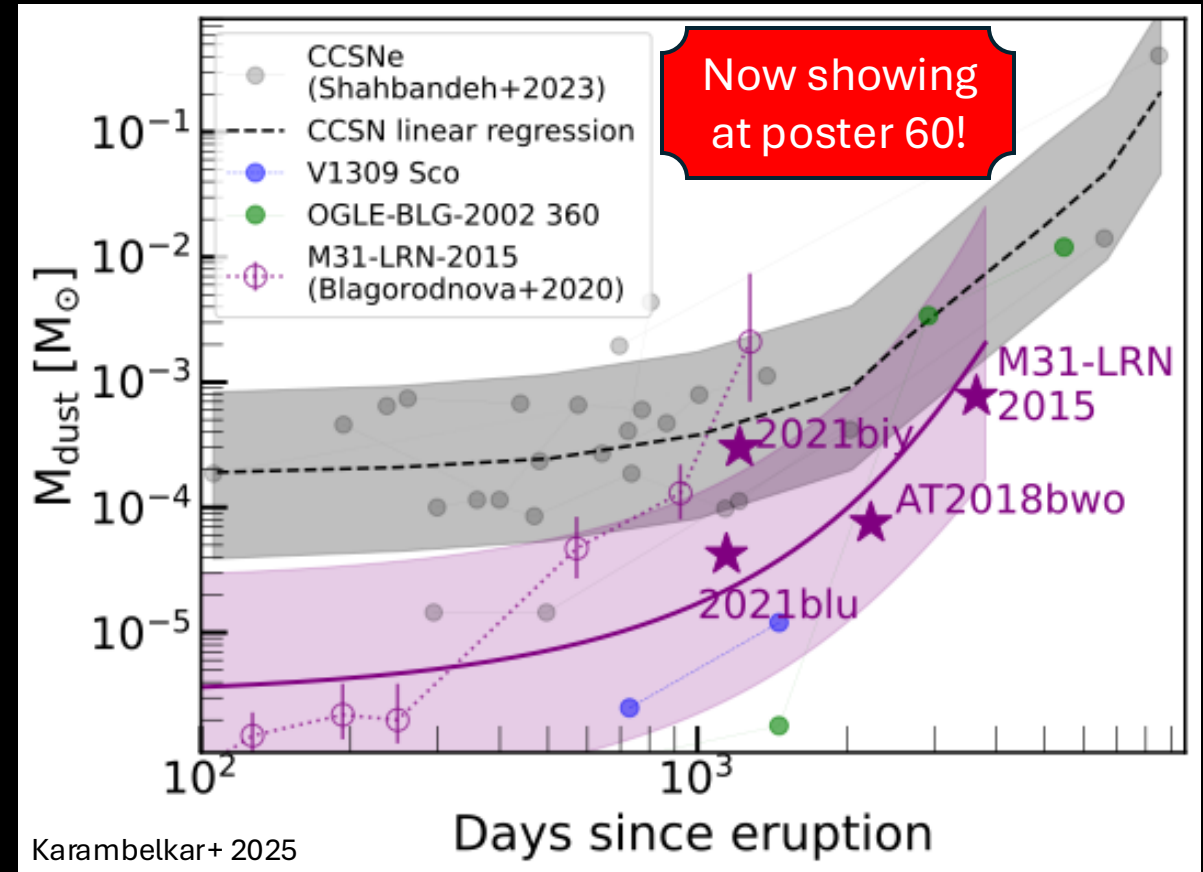
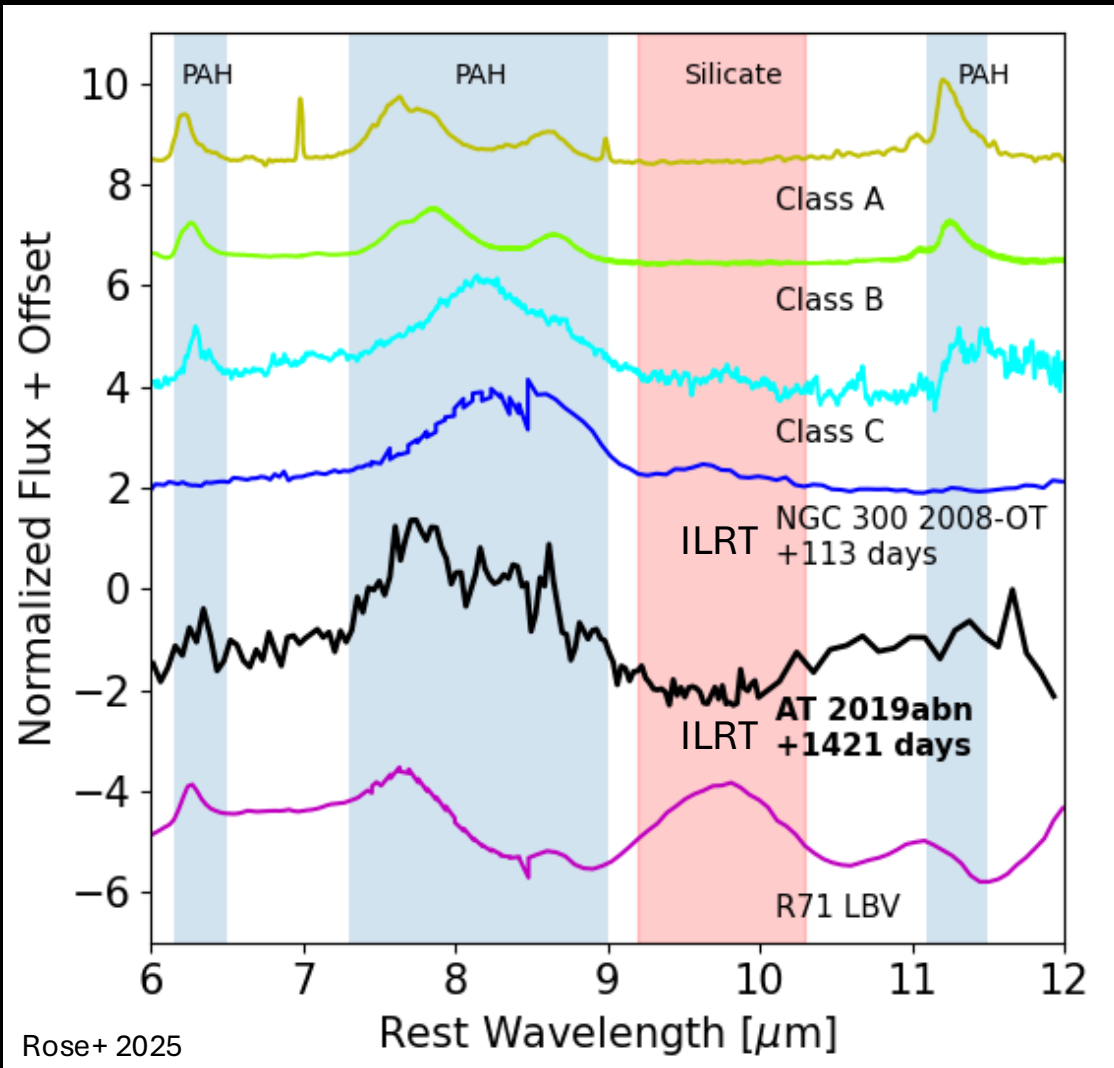


Image credit: R. Lau

Already evidence to suggest substantial dust contribution, including carbonaceous grains!



Dust produced by LRNe is $\sim 25\%$ dust produced by CCSNe!

Conclusions

- There are likely important contributors to the cosmic dust budget (especially at early times!) beyond AGB stars and core-collapse supernovae
- WR stars can produce significant amounts of carbonaceous dust at LMC-like metallicities
- The contributions to the cosmic dust budget of IR-bright “gap” transients remain underexplored
- Need IR missions for discovery and follow-up!

The future is IR-bright!

